



Event Horizon Telescope

Numerical Modeling of EHT Sgr A* Results

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with Ben Prather, Michi Baubock, George Wong, Vedant Dhruv, Abhishek Joshi
including work presented on behalf of the
Event Horizon Telescope Collaboration



Interferometry

millimeter VLBI



- 1.3mm VLBI network, $\Delta\theta \sim \lambda/D \sim (1.3\text{mm})/(2 R_{\oplus}) \sim 25\mu\text{as}$
- 2017 campaign: April 5, 6, 7, 10, 11; 6 targets, incl M87 & Sgr A*
- 8 telescopes at 6 sites

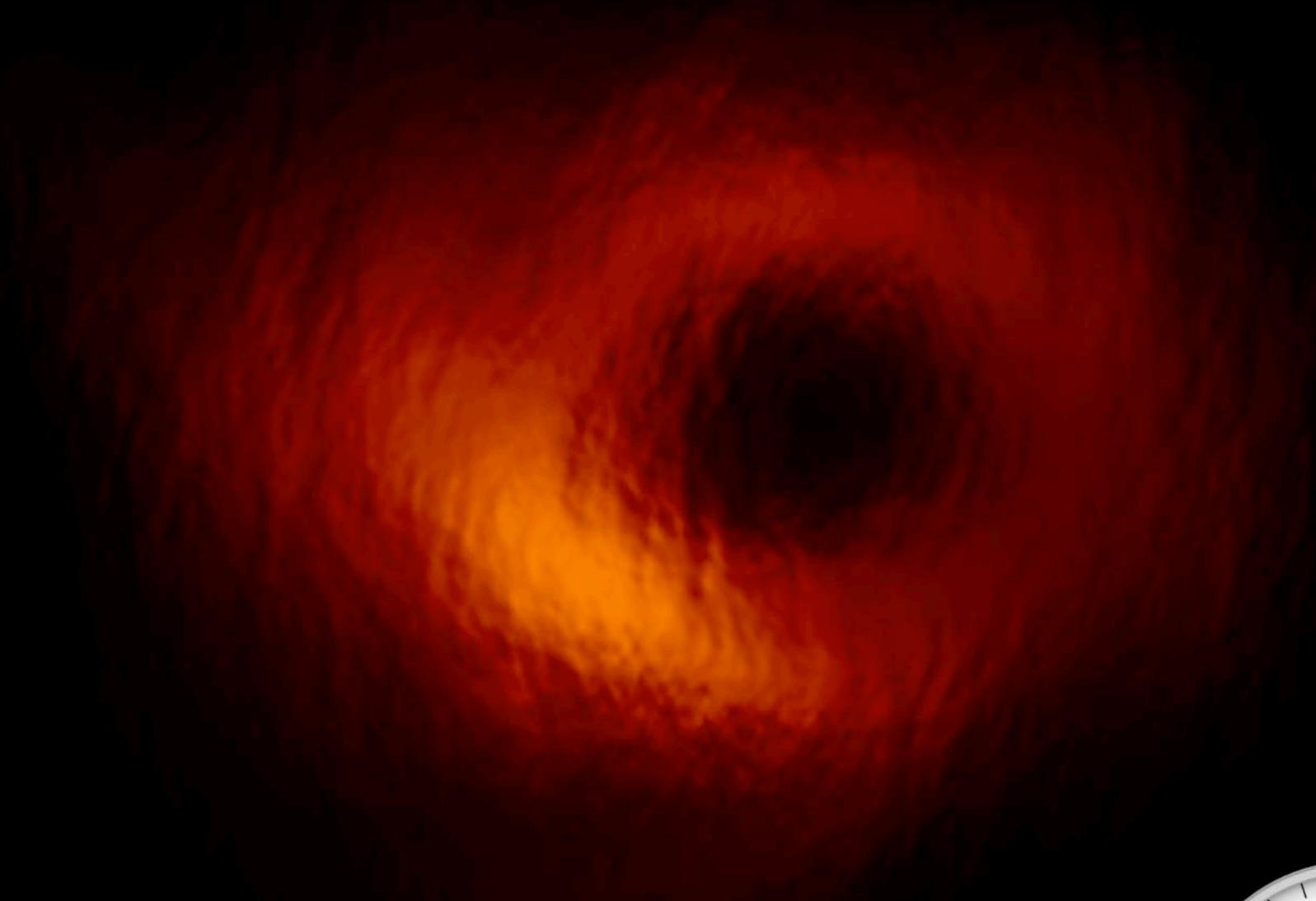


credit: ESO+

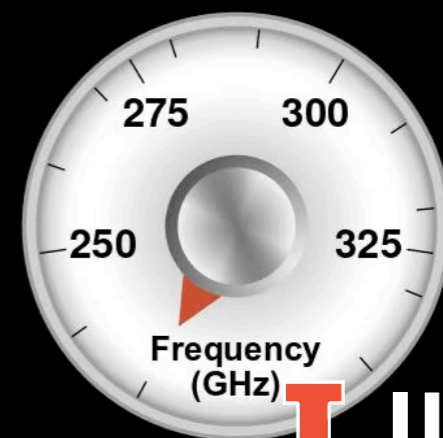
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Why is Sgr A* harder than M87*?



credit: M. Johnson



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Event Horizon Telescope



$$M \simeq 6.6 \times 10^9 M_{\odot}$$

$$\theta_g \equiv \frac{GM}{c^2 D} \simeq 3.8 \mu\text{as}$$

$$t_g \equiv \frac{GM}{c^3} \simeq 9.0 \text{ hr}$$

Sgr A*



$$M \simeq 4.1 \times 10^6 M_{\odot}$$

$$\theta_g \equiv \frac{GM}{c^2 D} \simeq 5.0 \mu\text{as}$$

$$t_g \equiv \frac{GM}{c^3} \simeq 20.4 \text{ sec}$$



M87



Sgr A*





credit: A. Joshi

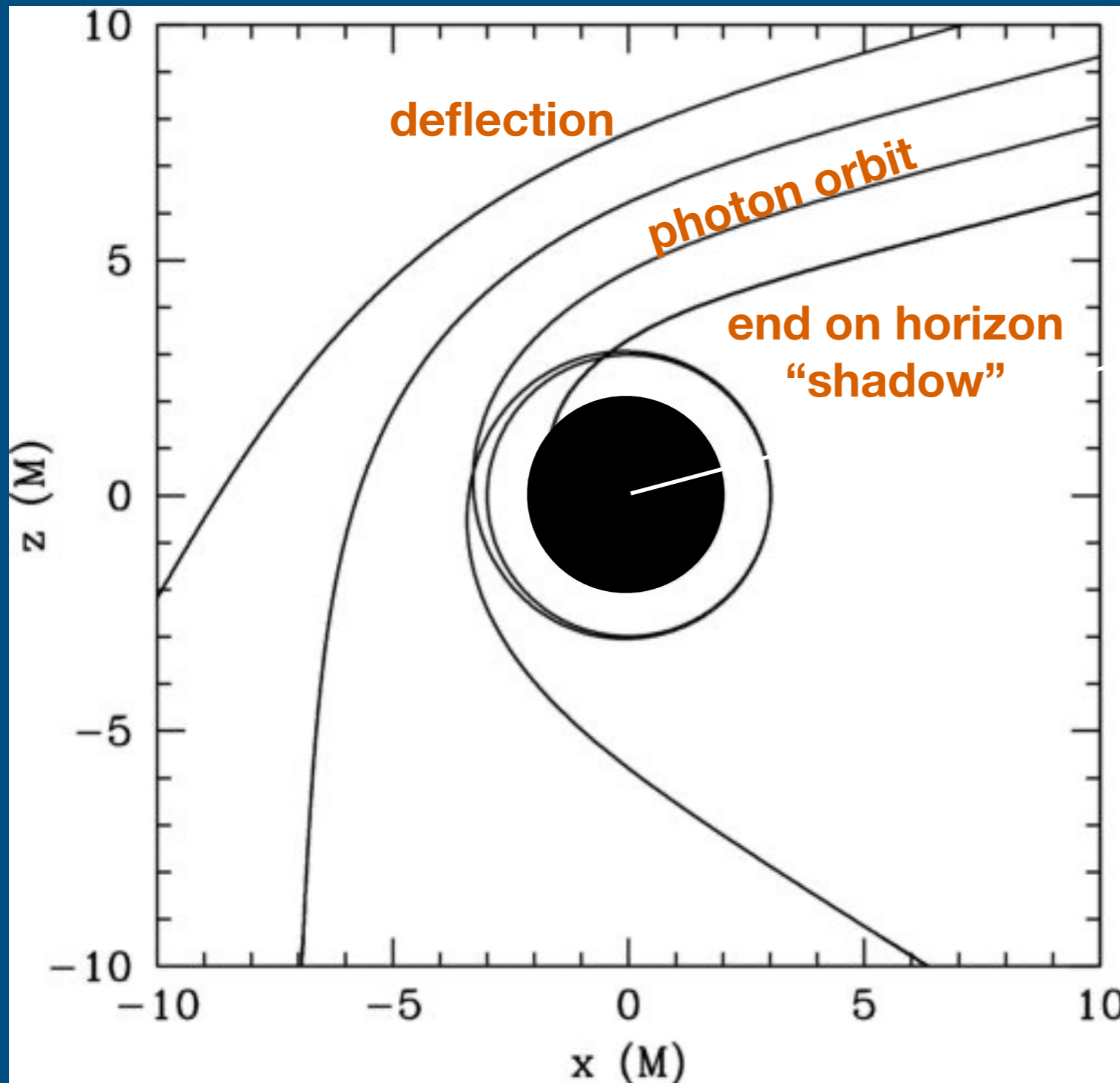
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Simulations and Numerical Experiments



credit: Raley et al., Uni Primary School

Gravitational Macrolensing



impact parameter
 $\sqrt{27}GM/c^2$



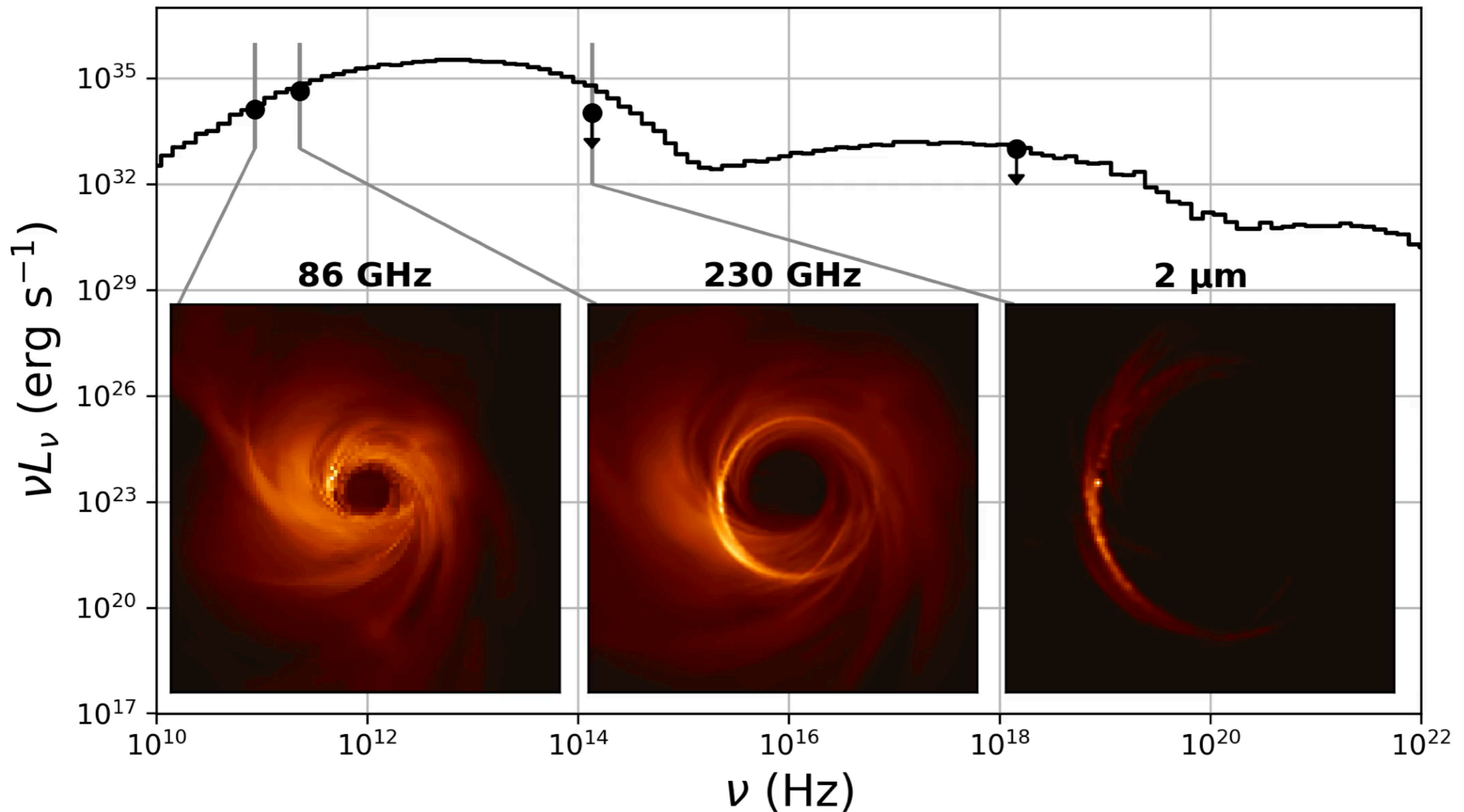
1. ideal GRMHD simulation
 2. assignment of electron DF
 3. radiative transfer calculation
- ⇒ Stokes $IQUV(v,x,y,t)$



credit: A. Joshi



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credit: B. Prather

Constraints on Models

EHTC et al. 2022, Paper I, V

Constraints

EHT-derived constraints

5, including 2nd moment, ring width

non-EHT constraints

4, including 86GHz size, flux

Variability constraints

2: structural variability, 230GHz flux var.

Model Parameters

1. spin
2. magnetization (MAD vs SANE)
3. inclination
4. electron DF assignment parameter R_{high}

credit: B. Prather



MAD

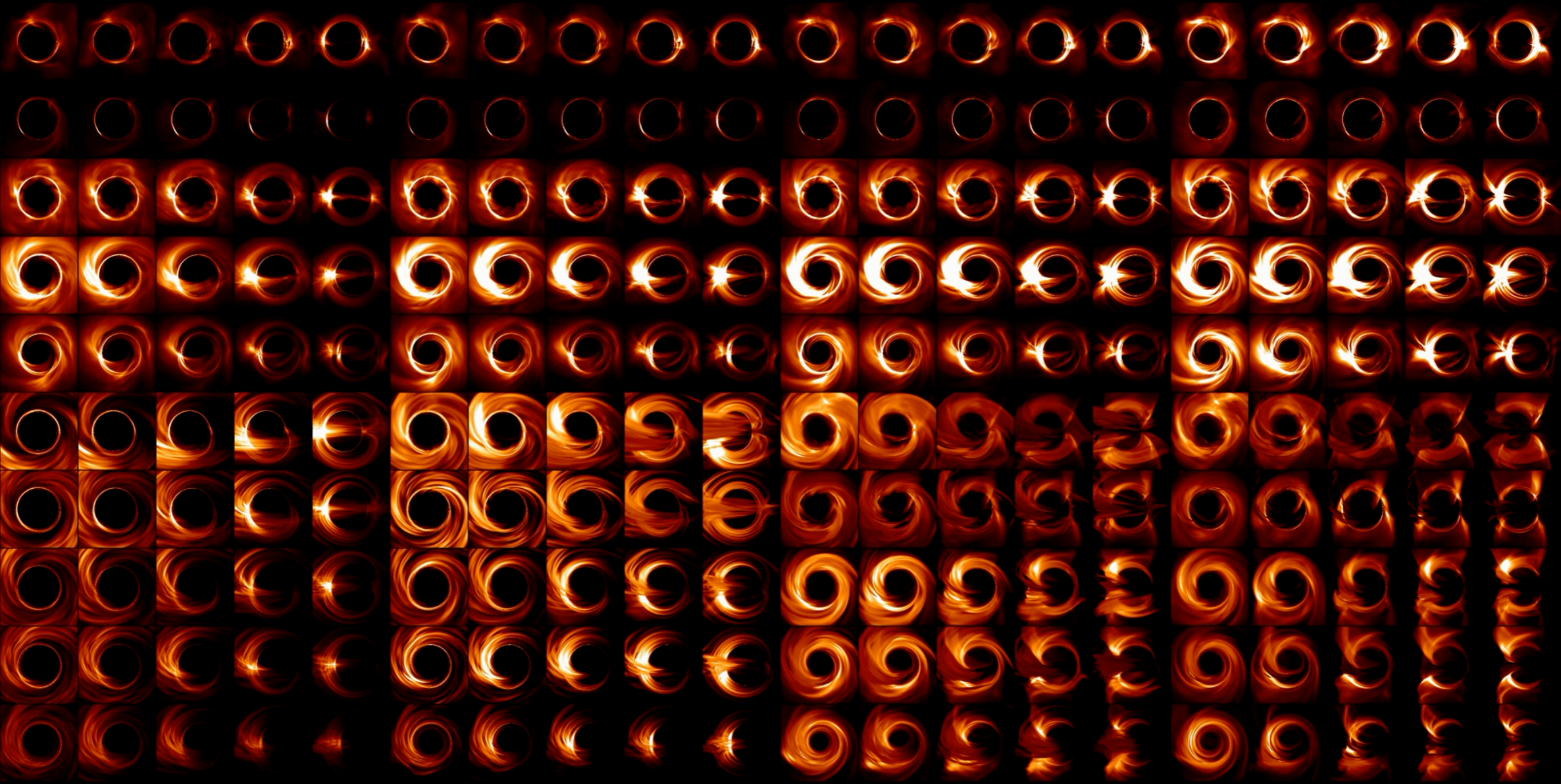
$$R_{\text{high}} = 160$$

$$a = +0.94$$

$$i = 90$$

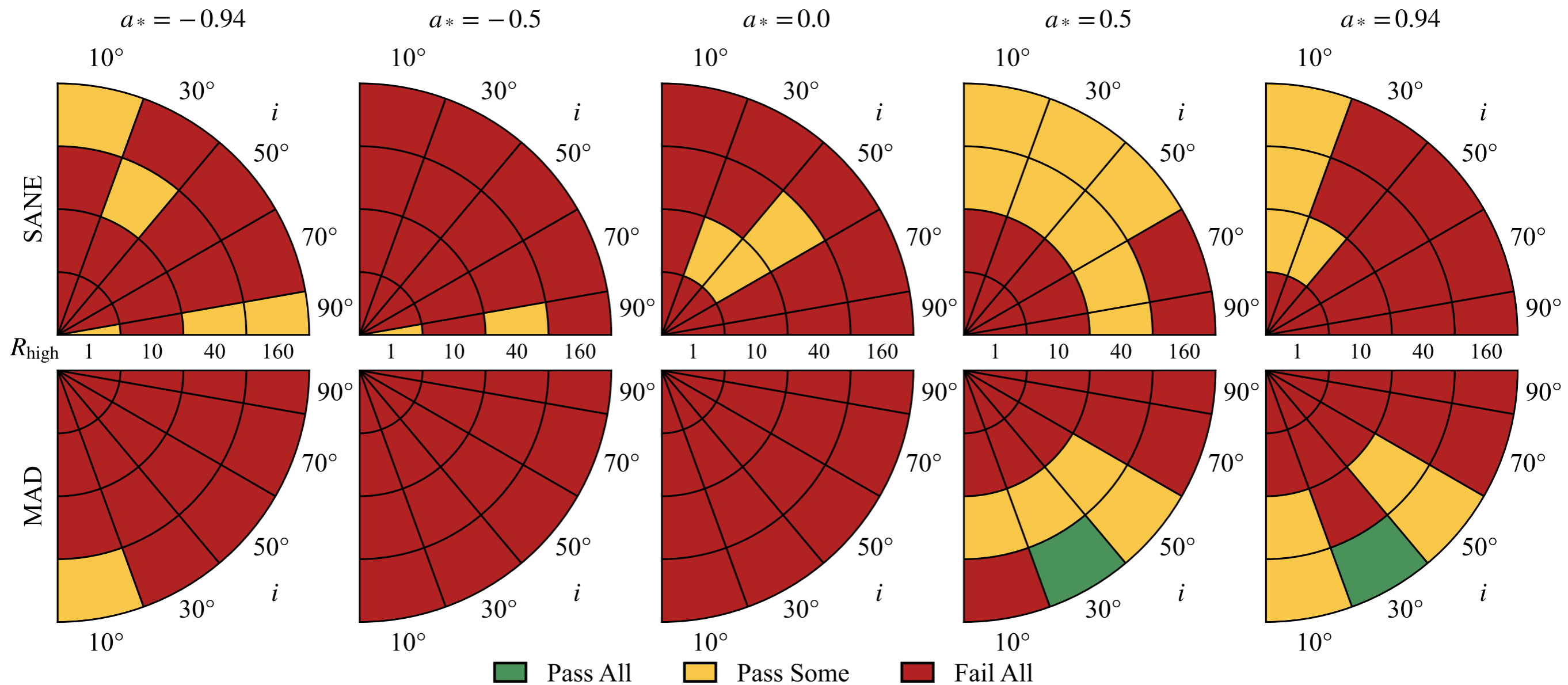
EHTC et al. 2022, Paper V

credit: B. Prather



EHTC et al. 2022, Paper V

All constraints except variability



EHTC et al. 2022, Paper I, V

Uncertainties

EHTC et al. 2022, Paper I, V

If you make a theory, for example, and advertise it, or put it out, then you must also put down all the facts that disagree with it, as well as those that agree with it.

-Feynman

- fluid model for collisionless plasma
 - $Kn \sim 10^5$
 - nonthermal electron DF
- boundary conditions (wind-fed?)
- model duration
- numerical resolution

Future Prospects

Future Prospects

polarization for Sgr A*

denser Fourier space coverage

movies

345GHz observations

space-based antenna?

more predictive numerical models including
viscosity and conduction

precision bothrology!

Blandford-Znajek Effect

NASA/ESA/Hubble Heritage

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Conclusions

- first image of galactic center black hole
- new, improved measurements on the way
- models explain all the data *except* variability
- variability crisis: origins of flow fluctuations?

