Searching for New Physics with Light (and Quantum technology)

Roni Harnik, **Fermilab Theoretical Physics Div. SQMS Center**

UIUC ICASU kickoff

PARTICLE PHYSICS

What are the basic degrees of freedom? What rules do they follow?

Accelarators, colliders, detectors, neutrino experiments, cosmic rays...

We are Curious!!!

What does the Universe contain? What is its history?

Telescopes, observatories, CMB, x-ray, gammaray, radio, direct detection experiments...

The Standard Model (of particle physics)

The Standard Model (of cosmology)

ENERGY DISTRIBUTION OF THE UNIVERSE

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Phierarchy problem

The Standard Model (of particle physics)

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ENERGY DISTRIBUTION OF THE UNIVERSE

The Standard Model (of particle physics)

Is there anything else? Could that be the dark matter? Can we learn more about cosmic history?

The Standard Model (of cosmology)

hierarchy problem

ENERGY DISTRIBUTION OF THE UNIVERSE

Overview - a non technical :

If HEP experimentalists and theorists are pursuing new approaches to address some of these question w/ technology being developed for HEP and for QIS (sQMs center).

- I New light particles that interact with photons: Dark Photons, Axions Advertising cavities (SQMS)
- D Ultra-high Q cavities and New physics searches: I Light shining through wall **O** Dark matter searches **Elight-by-light search**

D (Not new) light particles interacting with photons: Gravitational waves. O GW searches with cavities.

(and an advertisement for MAGIS-100 if there is time).

A good place to look for beyond standard mode is

The main actors:

Dark Photons

(both are good dark matter candidates, but are not assumed to be DM in most of this talk)

Much of the development in QIS technology involves manipulation of light, either in classical or quantum states.

New Particles Interacting w/ Light

Axions

Dark Photons - a Linear Extension

I Hypothesis: Add another photon to the rule book (and lets give it a mass)

O Why would such a particle exist?... Historically "copies" of particles did show up by surprise. (the muon: "who ordered that?!") O Even by surprise, a dark photon would teach us profound lessons. New

force of nature. Grand Unification, etc.

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Dark Photons - a Linear Extension

I Normal matter is not charged under the new photon. How will it interact?

quantum numbers can be in a superposition, "mix".

$$
|\psi\rangle = |\rho hoton\rangle
$$

I The dark Photon is a linear extension of QED with an effective Hamiltonian:

 $\mathcal{H} \supset \mathcal{H}_{QED} + \varepsilon \vec{E} \cdot \vec{E}' + \vec{B} \cdot \vec{B}'$

In quantum mechanics: two states which have the same

 $+ \varepsilon$ /dark photon)

 $\bigwedge\bigwedge\bigwedge\bigwedge$ \bigwedge \bigwedge

(and dark photon also has a mass)

I Photons are emitted by an EM current J, e.g oscillation dipole:

Graham et al, Phys.Rev.D 90 (2014)

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Transverse polarization \leftarrow

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I Photons are emitted by an EM current J, e.g oscillation dipole:

But now, the emitted state is $\langle \rho \rangle$ (photon) + ε (dark photon) !

Graham et al, Phys. Rev. D 90 (2014)

Transverse polarization $f \rightarrow$ Dark $f \mid u \times \infty$ ε^2

* in fact, ε here also depends on dark photon mass.

I Photons are emitted by an EM current J, e.g oscillation dipole:

But now, the emitted state $i\zeta^*$ | ρ

Reminiscent of Rabi (also neutrino) oscillations. Emitted state is a superposition of massless and massive state (mass difference \leftrightarrow detuning).

Graham et al, Phys. Rev. D 90 (2014)

$$
r\text{answer} \text{se polarization} \qquad \qquad \text{Dark flux } \propto \varepsilon^2
$$

$$
photon\rangle + \varepsilon
$$
 |dark photon\rangle !

* in fact, ε here also depends on dark photon mass.

|photon〉*+ ε |dark photon*〉

 \Box Recall a crucial difference b/w photons and dark counterpart - a mass.

 \Box The dark photon has a 3rd polarization!

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 $Dark flux \propto \epsilon^2$

 $\mathcal{J}(t)$

T Dark photon w/ logitudinal polarization

No oscillation, just emission of a feebly coupled state. In fact this flux is greater! *

Graham et al, Phys.Rev.D 90 (2014)

* in fact, ε here also depends on dark photon mass.

Dark Photon

Many constraints on the dark photon! (a review: Essig et al 1311.0029)

Axions - a Nonlinear Extension

I luvented to address a theoretical puzzle of the strong force: "Strong CP problem" (why dont (gluonic) E and B fields mix?)

I A nonlinear extension of QED, with a new invisible field:

$$
\mathscr{L} \supset \frac{a}{f} F^{\mu\nu} \tilde{F}_{\mu\nu} =
$$

I Axion phenomenology w/ background field is similar to dark photon. Mixing:

Pecci and Quinn (77)

-
-
- $\frac{a}{c} \vec{\epsilon} \cdot \vec{B}$ (1/f is a dimensionful coupling)
-

$$
\overrightarrow{B} \parallel \overrightarrow{J}(t)
$$

Photons polarized along a B field can mix with axions.

Axions and ALPs

Some technology -superconducting cavities

Superconducting Technology (and Fermilab)

□ SC technology is one of the most promising paths toward quantum computing with demonstrated results.

ISC qubits (frequencies of ~GHz) have achieved coherence times of ~ millisecond (Q~10⁶).

O Even higher Q systems were developed for accelerators

SRF cavities developed at Fermilab have achieved $Q \sim 10^{9-10}$.

Superconducting Technology (and Fermilab)

 \Box Also in the quantum regime:

1.3 No 5*.*² ⇥ ¹⁰¹⁰ ¹*.*⁰ ⇥ ¹⁰⁷ 0.17 In the TLS-dominated regime, and the TLS-dominated regime, and the TLS-dominated regime, and the TLS-dominated
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 $T_{\rm max}$ is $T_{\rm max}$ and $T_{\rm max}$ *f*⁰ (GHz) Oxide treatment *F*⁰ *F* ⁰ *Phys.Rev.Applied* 13 (2020) Romanenko et al.

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Can we use this as qubits? qudits?

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SQMS - Superconducting Quantum Materials and Systems

- \Box Fermilab's SQMS Center, funded by DOE's National Quantum Initiative, is leveraging this cutting edge SC research to quantum.
- **SQMS Will Reach New Limits of Coherence for Superconducting Qubits** \square Many institutions, including UIUC (Yoni Kahn).
	- SC and material science knowhow to improve regularity qubits. \mathbf{I}_{max} \ddot{a} -9
	- Use SRF cavities modes store quantum information, "3D architecture" 10 $\frac{1}{t}$

Leading US testbeds:

- *Google Sycamore*
- *IBM Hexagonal*
- *Rigetti Aspen*
- *Yale Single-mode*
- *UChicago Multi-mode*

Back to Dark photons, etc...

How can quantum sensing and technology help the search?

Lab Based Dark Photon Search

DRecall our radiating dipole:

Dark flux $\propto \epsilon^2$

"Light shining through wall (LSW)":

Jaekel et al (2006)

Can we sense this dark flux?

We know the frequency... (and the phase)!

Examples -Optical: ALPS, ALPSII, OSQAR **RF: CROWS**

O Consider two high quality cavities with with exactly same frequency

 $High Q \rightarrow we can store$ more photons

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 $High Q \rightarrow cavity can ring$ up for a longer time,

 $High Q \rightarrow we can store$ more photons

 $P_{\rm rec} \sim G^2 \, \epsilon^4 \left(\frac{m_{\gamma'}}{\omega}\right)^4 Q_{\rm rec} Q_{\rm em} P_{\rm em}$

O Consider two high quality cavities with with exactly same frequency

 $High Q \rightarrow cavity can ring$ up for a longer time,

Dark SRF the residual cross-talk, whereas for the medium power r

Proof of concept executed (w/ ingredients in the Fermilab pantry): a range of about the meaning of the same of the sa re i erroritud funit j/. $\begin{array}{ccc} \begin{array}{ccc} \mathbf{C} & \mathbf$ T CONCRPI

Plot by Zhen Liu (UMN)

\square Exciting future plans to exploit quantum sensing:

Preliminary DR setup, T~10 mK

Dark Photon Dark Matter

 \Box We are also testing hypothesis that a dark photon is the dark matter:

Dark matter

3

The dark matter is non relativistic Aphoton frequency is DM mass.

Need to scan cavity frequency

Dark Photon Dark Matter

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Dark matter

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The dark matter is non relativistic Aphoton frequency is DM mass.

Need to scan cavity frequency

Conceptually similar, but a background field is needed (recall axion is a nonlinear extension)

Axion searches

Axion Searches

 ω

 \mathcal{S}

Berlin et al (2019) $\sum_{i=1}^n$ with a total instrumented volume of $\mathcal{L}_{\mathcal{A}}$ axion and ALP data matter generated by the $\mathcal{L}_{\mathcal{A}}$

frequency = $m_a/2\pi$

Axion DM search - axion upconvert a pump field.

[GeV

 $\overline{}$ $\overline{}$

 \Box Our most devious scheme (Kahn and collaborators): searching for a nonlinear effect in as ingle cavity

Axion Induced Light-by-Light (and Euler-Heisenberg). Bogorad, Hook, Kahn_{uluc}, Soreq (2019)

| Light-by-light scattering among cavity modes because the e \sim 10 km s in the e \sim ton frequencies below the electron mass *m^e* = 511 keV

ight (and Euler-Heisenberg). Bogorad, Hook, Kahn_{uluc}, Soreq (2019)

! We are studying other sources of i mentally. $\ddot{}$ assumptions assumptions assumptions assumptions assumptions about the cavity $\ddot{}$ parameters, bandwidth, and measurement time; see text time; see text time; see text time; see text time; see t

Axion Searches

If cavity nonlinearities prove to

LSW with pump fields excited in both cavities.

Gao_{uluc}, RH (2021)

Gravitational wave searches

How are GWs related to axions?

Gravity waves - also a Nonlinear Extension

O GW's were only recently discovered. Extending the frequency window for GW detection is highly motivated

 \Box Gravity is also a nonlinear extension of QED:

and $GW's$ are fluctuations in the metric. $g = \eta + h$

 \Box Axion DM searches \leftrightarrow GW searches

 $\omega_q h B_0$

Understanding frame effects and (incorrect) existing literature was alot of fun.

Berlin et al [incliding Khanuluc, Schutte-Engeluluc] (2022)

-
- \mathscr{L})- \sqrt{g} $g_{\mu\sigma} g_{\nu\rho} F^{\mu\nu} F^{\sigma\rho}$
-

Note: correctly calculating the signal strength, is best done in the proper detector frame.

GHz Gravitational waves:

 \Box SQMS has potential to set new limits for GHz GW's (sources anyone...?)

Berlin et al [incliding Khan_{UIUC}, Schutte-Engel_{UIUC}] (2022)

Work in progress: Up-scattering experiment going to MHz frequencies!

Gravitational Waves with Matter Wave Interferometers

Jason Hogan of Stanford)

MAGIS-100: Gravitational waves with Atom Interferometry

njttjpo/ *Matter-wave Atomic Gradiometer Interferometric Sensor (MAGIS-100)* 7

 \Box 100 meters is a pathfinder toward a km scale gradiometer, as well as a space-based

Conclusions

D Axions and axion-like particles. **O** Dark photons

O Gravitational waves also couple to photons.

in the quest for new physics!

I New light particles that interact with photons are well motivated.

O New tools for enhanced coherence, control, and sensing of can be used

Deleted scenes

Dark photon and Axion searches with Quantum Optics

A new direction at the HEP-optics interface, in its infancy.

J. Estrada, RH, D. Rodriguez, M Senger - [arXiv:2012.04707](https://arxiv.org/abs/2012.04707), Accepted to PRX Quantum

Theory paper:

Nonlinear Optics with Bark

incoming light A Cartoon of optics:

presence of the other, every frigunity

<u>Dark SPDC:</u>

cently allowed the teleportunity of \mathcal{I} . $\mathcal{L}=\mathcal{L}^{\text{max}}$ imaging, or "interaction-free" image-free \mathcal{L}^{max} i tion about an object with the conditions. Note: the axion or dark photon have index of refraction of 1 (and a mass). dSPDC has significantly different phase matching conditions. Z **H**

Phase Matchi

D Two obse

[arXiv:2012.04707](https://arxiv.org/abs/2012.04707), Accepted to PRX Quantum

Two Hypotheses: New Particle search vs Dark Matter search

- production mechanisms.
- In the Wave-like DM category. Oscillating at $\omega = m_{DM}$.

• Axion-like particles, dark photons, B-L, are each dark matter candidates with nice

dark photons? axions?

dark photons? axions?

• Axion-like particles, dark photons, B-L, are each well motivated as mediators of

long range interactions that can be searched for.

Complementarity: New Particle search vs Dark Matter search

We produce our own dark stuff.

Light Shining through wall:

Complementary hypotheses. We want to test both!

Axions and ALPs

